REPORT DOC JENTATION PAGE

Form Approved OMB No. 0704-0188

Public reporting burden for this collection of information is estimated to average 1 hour per response, including the time for reviewing instructions, searching existing data sources, gathering and maintaining the data needed, and completing and reviewing this collection of information. Send comments regarding this burden estimate or any other aspect of this collection of information, including suggestions for reducing this burden to Department of Defense, Washington Headquarters Services, Directorate for Information Operations and Reports (0704-0188), 1215 Jefferson Davis Highway, Suite 1204, Arlington, VA 22202-4302. Respondents should be aware that notwithstanding any other provision of law, no person shall be subject to any penalty for failing to comply with a collection of information if it does not display a currently valid OMB control number. PLEASE DO NOT RETURN YOUR FORM TO THE ABOVE ADDRESS.

1. REPORT DATE (DD-	MM-YYYY)	2. REPORT TYPE			DATES COVERED (From - To)
15-12-2015		Final Technica	l Report		0/05/12 to 10/02/15
4. TITLE AND SUBTITLE				5a.	CONTRACT NUMBER
				BA	A-76-11-01
Impacts of Stratospheric dynamics on atmospheric behavior				r from 5b.	GRANT NUMBER
the ground to space solar minimum and sola			ar maximum	NO	0173-12-1G010
				5c.	PROGRAM ELEMENT NUMBER
6. AUTHOR(S)					PROJECT NUMBER
				76	-6077-12
Han-Li Liu				5e.	TASK NUMBER
				5f.	WORK UNIT NUMBER
7. PERFORMING ORGANIZATION NAME(S) AND ADDRESS(ES)					PERFORMING ORGANIZATION REPORT
					NUMBER
				20	11-027
University Corp					
Atmospheric Res					
3090 Center Gre					
Boulder, CO 803	305				
9. SPONSORING / MON	NITORING AGENCY	NAME(S) AND ADDRES	SS(ES)	10.	SPONSOR/MONITOR'S ACRONYM(S)
				HA	0
Naval Research	Laboratory			11.	SPONSOR/MONITOR'S REPORT
4555 Overlook A	_				
	00000				
Washington, DC	20375				
Washington, DC	20375				
Washington, DC	20375				NUMBER(S)
		MENT			NUMBER(S)
Washington, DC		MENT			NUMBER(S)
12. DISTRIBUTION / AV	/AILABILITY STATE		ation regults a		
12. DISTRIBUTION / AV	/AILABILITY STATE	MENT and its simula	ation results a		
12. DISTRIBUTION / AV	/AILABILITY STATE		ation results a		
12. DISTRIBUTION / AV	/AILABILITY STATE WACCM-X Model		ation results a		
12. DISTRIBUTION / AV	/AILABILITY STATE WACCM-X Model		ation results a		
12. DISTRIBUTION / AV	/AILABILITY STATE WACCM-X Model		ation results a		
12. DISTRIBUTION / AV Publications, V 13. SUPPLEMENTARY	/AILABILITY STATE WACCM-X Model		ation results a:		
12. DISTRIBUTION / AV Publications, V 13. SUPPLEMENTARY	/AILABILITY STATE WACCM-X Model NOTES	and its simula		re publicly	v available
12. DISTRIBUTION / AV Publications, V 13. SUPPLEMENTARY 14. ABSTRACT In this work, V	VAILABILITY STATE WACCM-X Model NOTES we have achie	and its simula	goal by (1) e	re publicly	available cesting, and delivering
12. DISTRIBUTION / AV Publications, V 13. SUPPLEMENTARY 14. ABSTRACT In this work, V WACCM-X model to	VAILABILITY STATE WACCM-X Model NOTES we have achies to the overal	and its simulated and its simu	goal by (1) en	re publicly nhancing, t	esting, and delivering team; (2) enabling coupling
12. DISTRIBUTION / AV Publications, V 13. SUPPLEMENTARY 14. ABSTRACT In this work, V WACCM-X model to WACCM-X with	VAILABILITY STATE WACCM-X Mode NOTES we have achie to the overal h the NAVDAS	eved the project	goal by (1) er or. F. Sassi and th the NRL SAMI:	nhancing, to the NRL to ionosphere	esting, and delivering team; (2) enabling coupling te and plasmasphere model;
12. DISTRIBUTION / AV Publications, V 13. SUPPLEMENTARY 14. ABSTRACT In this work, V WACCM-X model to of WACCM-X with (3) collaborate	VAILABILITY STATE WACCM-X Model NOTES we have achie to the overal h the NAVDAS ing on scient	eved the project of project PI, I system, and with	t goal by (1) en Or. F. Sassi and th the NRL SAMI E planetary wave	nhancing, to the NRL to sionospheres and their	testing, and delivering team; (2) enabling coupling te and plasmasphere model; in the lower
12. DISTRIBUTION / AV Publications, V 13. SUPPLEMENTARY 14. ABSTRACT In this work, V WACCM-X model to of WACCM-X with (3) collaborate	VAILABILITY STATE WACCM-X Model NOTES we have achie to the overal h the NAVDAS ing on scient	eved the project	t goal by (1) en Or. F. Sassi and th the NRL SAMI f planetary wave	nhancing, to the NRL to sionospheres and their	testing, and delivering team; (2) enabling coupling te and plasmasphere model; in the lower
12. DISTRIBUTION / AV Publications, V 13. SUPPLEMENTARY 14. ABSTRACT In this work, V WACCM-X model to of WACCM-X with (3) collaborate	VAILABILITY STATE WACCM-X Model NOTES we have achie to the overal h the NAVDAS ing on scient	eved the project of project PI, I system, and with	t goal by (1) en Or. F. Sassi and th the NRL SAMI f planetary wave	nhancing, to the NRL to sionospheres and their	testing, and delivering team; (2) enabling coupling te and plasmasphere model; in the lower
12. DISTRIBUTION / AV Publications, V 13. SUPPLEMENTARY 14. ABSTRACT In this work, v WACCM-X model to of WACCM-X with (3) collaboration thermosphere. A	VAILABILITY STATE WACCM-X Model NOTES we have achie to the overal h the NAVDAS ing on scient	eved the project of project PI, I system, and with	t goal by (1) en Or. F. Sassi and th the NRL SAMI f planetary wave	nhancing, to the NRL to sionospheres and their	testing, and delivering team; (2) enabling coupling te and plasmasphere model; in the lower
12. DISTRIBUTION / AV Publications, V 13. SUPPLEMENTARY 14. ABSTRACT In this work, V WACCM-X model to of WACCM-X with (3) collaborate	VAILABILITY STATE WACCM-X Model NOTES we have achie to the overal h the NAVDAS ing on scient	eved the project of project PI, I system, and with	t goal by (1) en Or. F. Sassi and th the NRL SAMI f planetary wave	nhancing, to the NRL to sionospheres and their	testing, and delivering team; (2) enabling coupling te and plasmasphere model; in the lower
12. DISTRIBUTION / AV Publications, V 13. SUPPLEMENTARY 14. ABSTRACT In this work, v WACCM-X model to of WACCM-X with (3) collaboration thermosphere. A	VAILABILITY STATE WACCM-X Model NOTES we have achie to the overal h the NAVDAS ing on scient	eved the project of project PI, I system, and with	t goal by (1) en Or. F. Sassi and th the NRL SAMI f planetary wave	nhancing, to the NRL to sionospheres and their	testing, and delivering team; (2) enabling coupling te and plasmasphere model; in the lower
12. DISTRIBUTION / AV Publications, V 13. SUPPLEMENTARY 14. ABSTRACT In this work, V WACCM-X model to of WACCM-X with (3) collaboration thermosphere. A 15. SUBJECT TERMS	WACCM-X Model WACCM-X Model NOTES We have achie to the overal h the NAVDAS ing on scient A final repor	eved the project of project PI, I system, and with	goal by (1) ender or F. Sassi and the NRL SAMI: f planetary wavese achievement.	nhancing, to the NRL to sionospheres and their sis attach	esting, and delivering team; (2) enabling coupling te and plasmasphere model; ir impact in the lower ned.
12. DISTRIBUTION / AV Publications, V 13. SUPPLEMENTARY 14. ABSTRACT In this work, v WACCM-X model to of WACCM-X with (3) collaboration thermosphere. A	WACCM-X Model WACCM-X Model NOTES We have achie to the overal h the NAVDAS ing on scient A final repor	eved the project of project PI, I system, and with	goal by (1) end of the NRL SAMI wave ese achievement.	nhancing, to the NRL to sionospheres and their is attack	testing, and delivering team; (2) enabling coupling re and plasmasphere model; ir impact in the lower ned.
12. DISTRIBUTION / AV Publications, V 13. SUPPLEMENTARY 14. ABSTRACT In this work, V WACCM-X model to of WACCM-X with (3) collaborat: thermosphere. A 15. SUBJECT TERMS 16. SECURITY CLASSI	WACCM-X Model WACCM-X Model NOTES We have achie to the overal h the NAVDAS ing on scient A final repor	eved the project of project PI, I system, and with tific studies of	goal by (1) ender or F. Sassi and the NRL SAMI: f planetary wavese achievement.	nhancing, to the NRL to sionospheres and their sis attach	testing, and delivering team; (2) enabling coupling re and plasmasphere model; ir impact in the lower ned.
12. DISTRIBUTION / AV Publications, V 13. SUPPLEMENTARY 14. ABSTRACT In this work, V WACCM-X model to of WACCM-X with (3) collaborate thermosphere. A 15. SUBJECT TERMS 16. SECURITY CLASSICAL REPORT	WACCM-X Model WACCM-X Model NOTES We have achie to the overal h the NAVDAS ing on scient A final repor	eved the project of project PI, I system, and with	goal by (1) end of the NRL SAMI wave ese achievement.	nhancing, to the NRL to sionospheres and their is attack	testing, and delivering team; (2) enabling coupling re and plasmasphere model; ir impact in the lower ned. 19a. NAME OF RESPONSIBLE PERSON Han-Li Liu 19b. TELEPHONE NUMBER (include area
12. DISTRIBUTION / AV Publications, V 13. SUPPLEMENTARY 14. ABSTRACT In this work, V WACCM-X model to of WACCM-X with (3) collaborat: thermosphere. A 15. SUBJECT TERMS 16. SECURITY CLASSI	WACCM-X Model WACCM-X Model NOTES We have achie to the overal h the NAVDAS ing on scient A final repor	eved the project of project PI, I system, and with tific studies of	goal by (1) end of the NRL SAMI wave ese achievement.	nhancing, to the NRL to sionospheres and their is attack	testing, and delivering team; (2) enabling coupling re and plasmasphere model; ir impact in the lower ned.

Standard Form 298 (Rev. 8-98) Prescribed by ANSI Std. Z39.18

20151229007

Final Report: Impacts of Stratospheric dynamics on atmospheric behavior from the ground to space under solar minimum and solar maximum conditions (Contract No.: N00173-12-1-G010 NRL)

Project Summary: Dynamical response to solar radiative forcing is a crucial and poorly understood mechanisms. We propose to study the impacts of large dynamical events on both the troposphere and the thermosphere during different phases of the solar cycle. The scientific objectives of this proposed research are intimately connected with the integrated response of the whole atmosphere to solar variability. In particular we compute and analyze the solar-induced variations of the following: (1) the penctration into the thermosphere of wave dynamics associated with disturbed events in the stratosphere; and (2) the influence of the stratosphere on the tropospheric climate during different phases of the solar cycle. In addition, a third objective of our research plan is to provide a fiducial simulation of the whole atmosphere up to 500 km which will allow the community to investigate in detail the sources and mechanisms that generate seasonal variations in the thermosphere (annual and semiannual variations). For this purpose, we will exercise the newly developed and updated extension of the Whole Atmosphere Community Climate Model (WACCM-X) to 500km which provides the most comprehensive ground-tothermosphere modeling capacity to date. To specify the stratospheric dynamical events as realistically as possible, the meteorology of the atmosphere below 90 km is constrained to the observed state using data assimilation products from the Naval Research Laboratory Atmospheric Variational Data Assimilation System (NAVDAS) or from the NASA Modern Era Retrospect Analysis for Research and Applications (MERRA). The quality of the model simulations (thus constrained) in the thermosphere will be assessed by comparing to the globally averaged mass density dataset developed at Naval Research Laboratory that covers the last 40 years and, where available, to composition, temperature and density profiles from the Global Ultraviolet Imager (GUVI) onboard of the NASA/TIMED satcllite.

Definition of NCAR's Role: NCAR PI H.-L. Liu will help the PI (F. Sassi) interfacing the DAS products with the WACCM-X for the different time periods and provide initial conditions for solar minimum and solar maximum. He will assist the rest of the team members in the validation of the thermospheric products.

Accomplishments

A. WACCM-X Development:

The NCAR Whole Atmosphere Community Climate Model with thermosphere and ionosphere extension (WACCM-X) has been further developed, tested, and delivered to the project PI, Dr. F. Sassi and the NRL team. Main developments include: (i) Developed and tested WACCM-X with specified dynamics capability (SD-WACCM-X). With this capability, the WACCM-X meteorological fields from the troposphere up to the mesopause can be specified by NAVDAS forecast results. This model setup (SD-WACCM-X/NAVDAS) has been used for case studies (discussed later). (ii) Developed a module to couple WACCM-X with an ionosphere/plasmasphere model. This module

allows the exchange of neutral fields from WACCM-X with ionospheric drifts, plasma density and electric conductivities from an ionosphere/plasmasphere model. Dr. F. Sassi and the NRL team has been able to couple WACCM-X with the NRL SAMI3 model. (iii) Implementation of time-dependent electron and ion temperature solver. In the middle and upper thermosphere, the thermal electron heating of the neutral atmosphere is the dominant heating source and thus important for thermosphere energetics. With this new implementation, the thermosphere temperature is higher and in better agreement with observations. (iv) Implementation of hydrogen escape flux at the model top boundary. A constant hydrogen flux was used in previous version of WACCM-X, which is partly responsible for the disagreement of various hydrogen species (e.g. H2O and CH4) with climatology. Recently an empirical formulation has been adapted from TIME-GCM and implemented in WACCM-X. This has resulted in better representation of these species by WACCM-X. Apart from these development efforts, we have also maintained and regularly update the WACCM-X to the latest CESM and CAM version.

B. Seientific Studies

B. 1 The Lower thermosphere during northern hemisphere winter of 2009

Numerical simulations were conducted using the SD-WACCM-X, constrained below 90 km by a combination of NASA's MERRA and the U.S. Navy's NOGAPS-ALPHA assimilation products. The period examined is January and February 2009, when a large stratospherie warming occurred on 24 January 2009, with anomalous circulation persisting for several weeks after the event. In this study, we focus on the dynamical response of the lower thermosphere up to 200 km. We find evidence of migrating and non-migrating tides, Rossby and Rossby-gravity modes, and Kelvin waves, whose amplitudes appear to be modulated at the times leading and following the stratospherie warming. While the Rossby, Rossby-gravity and Kelvin modes are rapidly dissipated in the lower thermosphere (above 110 km), the tides maintain substantial amplitude throughout the thermosphere, but their vertical structure becomes external above about 120-150 km. Most waves identified in the simulations decrease in amplitude in the thermosphere, indicating remote forcing from below and strong dissipation by molecular diffusion at high altitudes; however, the amplitude of the migrating DW1 tide increases in the thermosphere suggesting in situ foreing. We show that the amplitude of the tides (such as the DW1) changes as the background wind alters the vorticity in the tropics, which broadens or narrows the tropical waveguide. Our results also suggest that fast Rossby normal modes (periods ≤ 10 days) are excited by instability of the zonal mean wind distribution following the stratospheric warming.

Sassi, F., H.-L. Liu, J. Ma, and R. R. Garcia, The lower thermosphere during the northern hemisphere winter of 2009: A modeling study using high-altitude data assimilation products in WACCM-X, J. Geophys. Res., 118, doi:10.1002/jgrd.50632, 2013.

B.2 Westward traveling planetary wave events in the lower thermosphere during solar minimum conditions

The focus of this study is to describe how various dynamical conditions of boreal winter affect the dynamical behavior of the lower thermosphere (90-150 km). SD-

WACCM-X simulations were made whose dynamies is constrained by atmospheric specifications during recent and historical solar minimum conditions. The model simulations are earried out during solar minimum conditions and the results shown here diseuss the period January 1 - March 30 for five years (1995, 1996, 2008, 2009, and 2010). These years were selected because they include boreal winters without stratospherie warming (1995 and 1996), with modest or normal stratospherie warming (2008, 2010), and with a large and persistent stratospheric warming (2009). The ultimate goal of this study is to encapsulate the statistically significant dynamical behavior due to westward propagating, planetary-scale waves (wavenumber 1 and wavenumber 2) in the lower thermosphere that are associated with different stratospheric conditions. To this end we show that the westward zonal acceleration above about 80 km is by and large described by traveling waves with periods between 2 and 10 days. We show that the momentum carried by these waves is unlikely to affect directly the momentum budget of the extra-tropical lower thermosphere, where instead gravity-wave drag figures prominently. However, at the times leading to and following large stratospheric disturbances, we show prominent meridional propagation of wave activity from the midlatitudes toward the tropics. In eombination with strong eastward meridional wind shear, our results provide further evidence that such equatorward propagation of momentum in the lower thermosphere might influence the amplitude of equatorially trapped tides.

Sassi, F., and H.-L. Liu, Westward traveling planetary wave events in the lower thermosphere during solar minimum eonditions simulated by SD-WACCM-X, J. Atmos. Solar Terr. Phys., 119,11-26, doi: 10.1016/j.jastp.2014.06.009, 2014.

B.3 Traveling planetary-scale waves in the lower thermosphere: Effects on Neutral Density and Composition During Solar Minimum Conditions

The effects of breaking and dissipation of traveling, planetary scale Rossby waves (TPWs) in the lower thermosphere are investigated with respect to the mixing of neutral constituents, using WACCM-X whose meteorology below 92 km is eonstrained by atmospheric specifications obtained from operational weather forecast/data assimilation system. The simulations are carried out for the January-February months during the last solar minimum, 2008, 2009 and 2010. The Fourier spectra show that the amplitude of TPWs with periods between 3 and 10 days at mid-latitudes in the lower thermosphere are statistically significant in some years; the amplitude and phase of the band-pass filtered behavior is consistent with the behavior of the 5-day wave, the fundamental Rossby normal mode. A wavelet analysis using the S-transform (ST) shows that large variations with periods between 3 and 10 days ean occur in relatively narrow temporal windows (20–30 days) during boreal winter; by exploiting the eorrespondence between the timeaveraged ST and the Fourier speetrum, we determine the time dependent statistical significance of those amplitudes. The momentum flux entering the lower thermosphere during the times of TPW amplification is shown to be also large. Interestingly, these amplifications of the TPWs and the associated momentum flux in the thermosphere are not always associated with disturbed stratospheric events, such as stratospheric sudden warming (SSW). The zonal acceleration due to TPW events is largest at mid-latitudes in the lower thermosphere below 120 km, as expected, and in some winters it shows substantial magnitudes in the tropies during TPW events; the sub-tropical zonal

accelerations are consistent with Rossby wave encountering a surf zone at low latitudes, resulting in Rossby wave breaking and dissipation. The zonal acceleration is shown to be associated with a meridional diffusion, which is largest in the lower thermosphere where the zonal acceleration is also large. The ultimate effect on neutral density and composition is a meridional, down-gradient mixing; although this horizontal diffusion is largest below 110 km, the effects on the composition are amplified with increasing altitude, due to the diffusive separation of the thermosphere.

Sassi, F., H.-L. Liu, and J. T. Emmert, Traveling planetary-scale waves in the lower thermosphere: Effects on Neutral Density and Composition During Solar Minimum Conditions, J. Geophys. Res., under review.

B.4. The neutral dynamics during the 2009 sudden stratosphere warming simulated by different whole atmosphere models

This study compares simulations of the 2009 sudden stratospheric warming (SSW) from four different whole atmosphere models. The models included in the comparison are the Ground-to-topside model of Atmosphere and Ionosphere for Aeronomy, Hamburg Model of the Neutral and Ionized Atmosphere, Whole Atmosphere Model, and WACCM-X. The comparison focuses on the zonal mean, planetary wave, and tidal variability in the middle and upper atmosphere during the 2009 SSW. The model simulations are constrained in the lower atmosphere, and the simulated zonal mean and planetary wave variability is thus similar up to ~1 hPa (50 km). With the exception of WACCM-X, which is constrained up to 0.002 hPa (92 km) by NOGAPS-ALPHA, the models are unconstrained at higher altitudes leading to considerable divergence among the model simulations in the mesosphere and thermosphere. We attribute the differences at higher altitudes to be primarily due to different gravity wave drag parameterizations. In the mesosphere and lower thermosphere, we find both similarities and differences among the model simulated migrating and nonmigrating tides. The migrating diurnal tide (DW1) is similar in all of the model simulations. The model simulations reveal similar temporal evolution of the amplitude and phase of the migrating semidiurnal tide (SW2); however, the absolute SW2 amplitudes are significantly different. Through comparison of the zonal mean, planetary wave, and tidal variability during the 2009 SSW, the results of the present study provide insight into aspects of the middle and upper atmosphere variability that are considered to be robust features, as well as aspects that should be considered with significant uncertainty.

Pedatella, N. M., T. Fuller-Rowell, H. Wang, H. Jin, Y. Miyoshi, H. Fujiwara, H. Shinagawa, H.-L. Liu, F. Sassi, H. Schmidt, V. Matthias, and L. Goncharenko, The neutral dynamics during the 2009 sudden stratosphere warming simulated by different whole atmosphere models, J. Geophys. Res., 119, 1306-1324, doi:10.1002/2013JA019421, 2014.

B.5. Ionosphere variability during the 2009 SSW: Influence of the lunar semidiurnal tide and mechanisms producing electron density variability

WACCM-X, eonstrained by NOGAPS-ALPHA, and TIME-GCM are used to investigate ionosphere variability during the 2009 sudden stratosphere warming (SSW). The simulations reveal notable enhancements in both the migrating semidiurnal solar (SW2) and lunar (M_2) tides during the SSW. The SW2 and M_2 amplitudes reach $\sim 50 \text{ m s}^{-1}$ and $\sim 40 \text{ m s}^{-1}$, respectively, in zonal wind at E region altitudes. The dramatic increase in the M_2 at these altitudes influences the dynamo generation of electric fields, and the importance of the M_2 on the ionosphere variability during the 2009 SSW is demonstrated by comparing simulations with and without the M_2 . TIME-GCM simulations that incorporate the M_2 are found to be in good agreement with Jieamarca Incoherent Scatter Radar vertical plasma drifts and Constellation Observing System for Meteorology, Ionosphere, and Climate (COSMIC) observations of the maximum F region electron density. The agreement with observations is worse if the M_2 is not included in the simulation, demonstrating that the lunar tide is an important contributor to the ionosphere variability during the 2009 SSW. We additionally investigate sources of

the F region electron density variability during the SSW. The primary driver of the electron density variability is changes in electric fields. Changes in meridional neutral winds and thermosphere composition are found to also contribute to the electron density variability during the 2009 SSW. The electron density variability for the 2009 SSW is

therefore not solely due to variability in electric fields as previously thought.

Pedatella, N. M., H.-L. Liu, F. Sassi, J. Lei, J. L. Chau, and X. Zhang, Ionosphere variability during the 2009 SSW: Influence of the lunar semidiurnal tide and mechanisms producing electron density variability, J. Geophys. Res., 119, doi:10.1002/2014JA019849, 2014.

B.6. Day-to-day variation of the equatorial electrojet during quiet periods

It has been known for a long time that the equatorial electrojet varies from day to day even when solar and geomagnetic activities are very low. The quiet time day-to-day variation is eonsidered to be due to irregular variability of the neutral wind, but little is known about how variable winds drive the electrojet variability. We employ a numerical model introduced by Liu et al. (2013), which takes into account weather changes in the lower atmosphere and thus ean reproduce ionospherie variability due to forcing from below. The simulation is run for May and June 2009. Constant solar and magnetospheric energy inputs are used so that day-to-day changes will arise only from lower atmospherie foreing. The simulated electrojet current shows day-to-day variability of $\pm 25\%$, which produces day-to-day variations in ground level geomagnetic perturbations near the magnetic equator. The eurrent system associated with the day-to-day variation of the equatorial electrojet is traced based on a eovariance analysis. The current pattern reveals return flow at both sides of the electrojet, in agreement with those inferred from groundbased magnetometer data in previous studies. The day-to-day variation in the electrojet eurrent is compared with those in the neutral wind at various altitudes, latitudes, and longitudes. It is found that the electrojet variability is dominated by the zonal wind at 100–120 km altitudes near the magnetic equator. These results suggest that the response of the zonal polarization electric field to variable zonal winds is the main source of the day-to-day variation of the equatorial electrojet during quiet periods.

Yamazaki, Y., A. D. Richmond, A. Maute, H.-L. Liu, N. M. Pedatella and F. Sassi, On the day-to-day variation of the equatorial electrojet during quiet periods, J. Geophys. Res., 119, doi:10.1002/2014JA020243, 2014.